A timeliness to potentials.

Lesson Two.





In our last lesson we encountered the metric g=gijdxi @ dxj: V > V*, which took vectors to their duals, 1-forms. Now we will consider a specific metric M, The Minkowski metric. M = Mun dx " & dx carries cineek indices denoting (t, x, y, z) and describes a space time geometry! n is defined by the mapping:

$$\partial_t \xrightarrow{\eta} - dt$$
, $\partial_x \xrightarrow{\eta} dx$,
 $\partial_y \xrightarrow{\eta} dy$, $\partial_z \xrightarrow{\eta} dz$.

Let's explore the metric of by computing the magnitudes of a couple of vectors. if $\vec{\sigma} = \lambda \partial_x$: $\underline{if} \quad \overline{u} = \partial \partial_{\mathbf{f}};$ ⇒ | טֿ|² = ଓ(טֿ) = ๆ טֿטֿ $\Rightarrow |\vec{u}|^2 = \vec{u}(\vec{u}) = \eta \vec{u} \vec{u}$ and $\vec{v} = 2 d d d$ and $\tilde{u} = -\lambda dt$ $\Rightarrow |\vec{v}|^2 = 2 dx (2 \partial_x) = 4$ $\Rightarrow |\hat{u}|^2 = -\lambda dt (\lambda \partial_t) = -4$ JJ dr <u>_ū_</u> dt X We can see here $dx(\vec{v}) = \lambda$. We can see here $-dt(\overline{u}) = -2$.

The vector it has a negative squared norm! it is an example of a timelike vector. As $d(\vec{u})$ is a scalar, we know \vec{u} is a timelike vector in all coordinate systems. Let us now compute the norm of a vector $\mathbf{\vec{l}} = \partial_{\mathbf{t}} + \partial_{\mathbf{x}}$. $|\vec{l}|^2 = \vec{l}(\vec{l}) = \eta \vec{l} \vec{l} = \eta (\partial_t + \partial_x) (\partial_t + \partial_x)$ $= (-dt + dx)(\partial_t + \partial_x)$ $= -dt \partial_t - dt \partial_x + dx \partial_t + dx \partial_x = 0.$ $\Rightarrow \overline{l}$ is an example of a null vector! Let's see how 1-form addition can help us understand null vectors.

Here we describe now to add 1-forms-togethes. $\mathbf{\Lambda}$ A+B 7 A B-A Y We consider the unit: There are two ways to intersect both A and B, =+2 and =0. The unit that fulfills both is: See key Note that this is with A and B having the orientations \uparrow and \rightarrow as given. 1-forms have a sense of direction just as vectors do: $\leftarrow = -1 \times \leftarrow \leftarrow$.



dt, dx, and dx+dt are depicted nese. We can also draw -dt+dx, by connecting the diagonals across the units in the other direction. We can see now how $\tilde{\mathcal{I}}(\tilde{\mathcal{I}}) = 0$, they don't intersect at all! We summarize: $|\vec{U}|^2 > 0 \Rightarrow$ space like $|\vec{u}|^2 < 0 \Rightarrow$ timelike $|\vec{v}|^2 = 0 \Rightarrow$ lightlike

Two coordinate systems with metric 1, one Moving at velocity $\vec{U} = U \hat{x}$ relative to the other, are related by a Lorentz transformation: $\Lambda: t' = \chi t - \chi \beta x y' = y$ $\chi' = \chi x - \chi \beta t z' = z', \Lambda u' = \begin{pmatrix} \chi - \chi \beta z \\ -\chi \beta z \\ -\chi \beta z \end{pmatrix}$ where $\beta = U/C$, C = speed of light, and we have $\gamma = (1 - \beta^2)^{-1/2}$ chosen units where C = 1. (such a / is sometimes called a hyperbolic rotation.) Let us consider the vector $p^{\mu} = m\partial_{\xi}$, with m being a mass term. what is put in the moving frame?

 $P^{M} = \frac{\partial \chi^{M}}{\partial \chi^{M}} P^{M} = \Lambda^{M'}_{\mu} P^{M}$ $\Rightarrow \rho^{\mu} \partial_{\mu'} = \gamma m \partial_{t} - \gamma \beta m \partial_{x}.$ if $\mathbf{U} << c$, $\gamma = (1 - \frac{\mathbf{U}^2}{c^2})^{-1/2} \approx 1 + \frac{1}{2} \frac{\mathbf{U}^2}{c^2}$. $\Rightarrow P^{\mu} \partial_{\mu'} \approx (m + \frac{1}{2} m \frac{\sigma^2}{c^2}) \partial_{\epsilon} - (m \frac{\sigma}{c} + \frac{1}{2} m \frac{\sigma^3}{c^3}) \partial_{\kappa}$ → p^mC²∂_µ × mC² + ±mv² ∂_t - mvC ∂_x restmass The non-relativistic energy The non-relativistic kinetic energy! momentumi We see that a particle's spatial momentum arises from its motion through time being (nyperbolically) rotated into a spatial dimension!

Let's do a relavistic example. t = 4/5 c, $\chi = 5/3 \Rightarrow$ if $q_{M}' = m \partial_{\epsilon}r$, $p_{M} = m \partial_{\epsilon}$, q_{t} $q_{M}' = \chi m \partial_{\epsilon} + \chi \beta m \partial_{\chi}$. $q_{M}' = 5/3 m \partial_{\epsilon} + 4/3 m \partial_{\chi}$.

pⁿ and qⁿ are examples of 4-momentum vectors, each seen at rest in their own frame, both seen at rest in their own frame, whose time component gives rise to spatial momentum (p^{xr}, q^x) when seen in a moving frame. What other physical quantities might behave like this?

Having developed an intuition for now vectors such as pully and que du transform, we will also look at now 1-forms such as A = Andx" transform, and how to plot their components. Suppose we have A'= - 4'dt'. what is A, the 1-form in the rest frame? $(U = 4/5C, \chi = 5/3)$. $A_{\mu} = \Lambda_{\mu}^{\mu} A_{\mu} \Rightarrow A_{\mu} dx^{\mu} = -\chi \phi' dt + \chi \beta \phi' dx$ \Rightarrow Andx" = -5/36 dt + 4/36 dx.

what is this object A?

 $A_{\mu} = g_{\mu\nu} A^{\nu}$ is the dual to $A^{\mu} = (\phi, \bar{A})$, where ϕ is the scalar potential and \bar{A} is the vector potential! They fulfill the following equations:

$$\vec{E} = -\vec{\nabla}\phi - \partial_e \vec{A}$$
$$\vec{B} = \vec{\nabla} \times \vec{A}$$

In the relativistic formulation of electrochynamics the vector potential \hat{A} is promoted to the <u>four-potential</u> A^{μ} by noting that the scalar potential ϕ is actually the time component A^{\pm} of the four-potential.

With Andrew' = - &' dt' and Andx" = -5/36 dt+4/36 dx, and additionally assuming that $\phi' = 1$, we are ready to plot A in both coordinate systems. with dx $(\vartheta_v) = \M_v , we have $-\$/3dt(\vartheta_t) = -\$/3$ and -5/3 dt (30) =-5 => 30e traverses -5 units demarcated by -5t3 dt. Likewise, 4t3 dx (37) = 4. -5/3dt, 4/3dx we can represent these relations with: $\partial_t = \frac{1}{2} + \frac{1}{$

Why do we want to describe the four-potential A as a 1-form? Consider the equation: $B = \nabla x A$. We can integrate both sides over a 2D manifold M. $\iint_{\mathbf{H}} \vec{\mathbf{B}} \cdot d\vec{\mathbf{a}} = \iint_{\mathbf{H}} \vec{\nabla} \times \vec{\mathbf{A}} \cdot d\vec{\mathbf{a}}$ we can then apply Stokes' theorem: $\iint \vec{B} \cdot d\vec{a} = \oint \vec{A} \cdot d\vec{k}$ We notice that $\vec{A} \cdot d\vec{k} = A_i dx^i$, a 1-form! So we will be interested in evaluating integrals of the form $\int A_i dx^u = \int A_i dx^i = \int A_i d$

The second equation to consider is the equation: $\vec{E} = -\vec{\nabla} \vec{\phi} - \partial_{\vec{E}} \vec{A}$ Looking at $\vec{\nabla}\phi$, $\vec{\nabla}\phi = \frac{\partial\phi}{\partial x} + \frac{\partial\phi}{\partial y} + \frac{\partial\phi}{\partial z} + \frac{\partial\phi}{\partial z}$. we find that The dual to $\vec{\nabla}_{\phi}$ with the flat space (g=1) metric is: $\mathbf{1} \overline{\nabla} \phi = \partial \phi dx + \partial \phi dy + \partial \phi dz = \partial i \phi dz^{i}$ By the chain rule, this is just the differential du, which is a 1-form, depicted as level curves of 6!

Finally, let's discuss the term
$$\partial_{\varepsilon} \hat{A}$$
.
In index notation, this is $\frac{\partial A^{i}}{\partial \chi t}$ (where $\chi t \equiv \varepsilon$).
We also have $\vec{\nabla} \phi$, which is $\frac{\partial A^{t}}{\partial \chi^{i}}$. Then, we have
 $E^{i} = -\frac{\partial A^{t}}{\partial \chi^{i}} - \frac{\partial A^{i}}{\partial \chi t} = -\partial_{i}A^{t} - \partial_{\varepsilon}A^{i}$.
 $B^{i} = \frac{\partial A^{k}}{\partial \chi^{i}} - \frac{\partial A^{i}}{\partial \chi t} = -\partial_{i}A^{t} - \partial_{\varepsilon}A^{i}$.

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(Note that in flat space upper and lower indices carry no additional meaning.)

The equations $E^{\hat{\nu}} = -\partial_{\hat{\nu}}A^{\hat{\tau}} - \partial_{\hat{\nu}}A^{\hat{\nu}}$ and $B^{\hat{\nu}} = \partial_{\hat{\nu}}A^{\hat{\nu}} - \partial_{k}A^{\hat{\nu}}$ are unified in relativistic electrodynamics by the following equation:

$$F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}$$

This is the electromagnetic field strength tensor, the Faraday tensor, and it is a 2-form!

In the next lesson, we will see how to express F in terms of basis 2-forms. The potential equations $(\vec{E} = -\vec{\nabla}\phi - \partial_E\vec{A}, \vec{B} = \vec{\nabla}\times\vec{A})$ have introduced us to three differential forms: $d\phi = \partial_{\mu}\phi dx^{\mu}$, a = 1-formA = Andra, a l-form F, a 2-form. They are called so because they can be placed underneath an integral sign and integrated, as we saw with J.A. How do we visualize these?



A = Andra we will visualize the same way, with one important difference: the curves depicting A will not necessarily close! We will see this next time.

Summary: The Minkowski metric of highlights now vectors and 1-forms (and their depictions!) are both needed to describe the geometry of spacetime. We also began to see how electromagnetism could be described using differential forms, something we first explored by considering the timeliness of potentials.